## METR 5123, Advanced Atmospheric Dynamics II Alan Shapiro, Instructor

## **REVIEW OF SURFACE GRAVITY WAVES**

First, we'll review some notation and concepts that will be applicable to many kinds of waves (not just surface gravity waves).

Often we'll work w/ waves of the form:

some flow property ~ 
$$a \sin\left[\frac{2\pi}{\lambda}(x-ct)\right]$$
 or  $a \cos\left[\frac{2\pi}{\lambda}(x-ct)\right]$ 

a is amplitude

 $\frac{2\pi}{\lambda}(x - ct)$  is <u>phase</u> of wave

 $\lambda$  is <u>wavelength</u>

c is phase speed

Why consider waves of this form?

-- Many waves in atmosphere "look" like sines or cosines.

-- Many waves approximately satisfy <u>linear const coeff odes and</u> <u>pdes</u> that permit sin and cos solns. Can use <u>Fourier analysis</u> to get solns of complicated problems by summing sin and cos solns of various amplitudes and wavelengths.

A closer look at wave parameters:

phase: 
$$\frac{2\pi}{\lambda}(x - ct)$$

Wave repeats itself when <u>phase changes by  $2\pi$ </u>. So, at a fixed moment in time, phase changes by  $2\pi$  when x changes by  $\lambda$ . Hence the name wavelength for  $\lambda$ .

$$k \equiv \frac{2\pi}{\lambda}$$
 is wavenumber, # waves in (dimensional) length of  $2\pi$ .

e.g., if  $\lambda = \pi$  (meters) then there are 2 waves in  $2\pi$  meters.  $k = \frac{2\pi}{\pi m} = 2 m^{-1}$ .

At a fixed point, phase changes by  $2\pi$  when time changes by  $\frac{\lambda}{c}$ So wave <u>period</u> is:  $T \equiv \frac{\lambda}{c}$ .

$$v \equiv \frac{1}{T}$$
 is frequency, # of oscillations per unit time

 $\omega \equiv \frac{2\pi}{T} = 2\pi v$  is circular (or radian) frequency.

Since  $T = \frac{\lambda}{c} \rightarrow \omega = 2\pi \frac{c}{\lambda} = kc$ 

 $\therefore$  c =  $\frac{\omega}{k}$ 

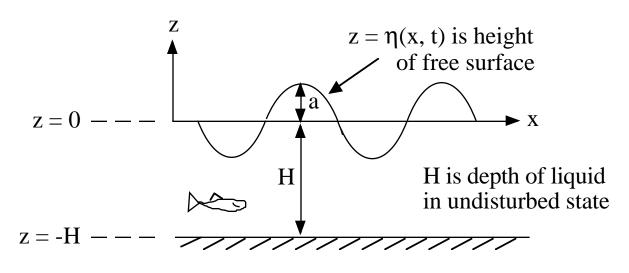
alternate expressions:

$$sin\frac{2\pi}{\lambda} (x - ct)$$
$$sin[k(x - ct)]$$
$$sin(kx - \omega t)$$

Motion of crests/troughts is motion of a geometric pattern. <u>Fluid</u> does not generally move with the wave pattern, e.g. Rossby waves, e.g., ocean waves might propagate toward shore at 20 m/s, but fluid moves toward and away from shore at  $\sim 1 \text{ m/s}$ .

## Surface gravity waves

Consider 2-D motion in the x-z plane of a liquid w/ a free surface (e.g. air/sea interface) [a pool of cold air underlying warm air is similar but more complicated]



-- assume disturbance is "infinitesimal", i.e., a is so small that:

a << H (a/H << 1)

and  $a \ll \lambda$   $(a/\lambda \ll 1)$ 

Because of this, nonlinear accel<sup>n</sup> terms are very small compared to local derivs (products of small quantities are really small)

-- neglect friction

-- assume T << rotation period of earth, or equivalently  $\omega >> f$  $\therefore$  can safely neglect Coriolis force.

-- assume fluid was initially at rest (so irrotational) and waves created irrotationally. So from Kelvin's Th<sup>m</sup>, the flow <u>will</u> always be irrot. So  $\vec{\omega}(t) = 0$ 

 $\therefore \quad \vec{u}(t) = \nabla \phi(t)$ 

-- assume flow is incompressible,  $\nabla \cdot \vec{u} = 0$ 

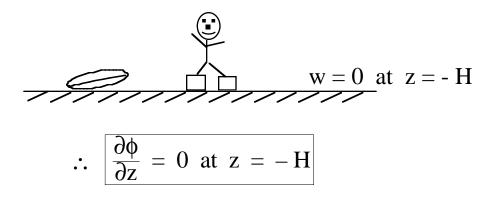
Substituting in  $\vec{u} = \nabla \phi$  we get:

 $\nabla\cdot\nabla\varphi\ =\ 0$ 

 $\therefore \nabla^2 \phi = 0 \text{ Laplace's eqn.}$ 

$$\frac{\partial^2 \phi}{\partial x^2} + \frac{\partial^2 \phi}{\partial z^2} = 0$$

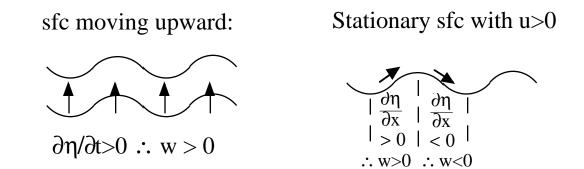
Need b.c.s to solve it. At bottom: <u>impermeability cond</u>n:



On top (free surface): a kinematic b.c. and a dynamic b.c.

<u>Top Kinematic b.c.</u>: A fluid element on free sfc remains on that sfc no matter how sfc moves or deforms.  $\therefore z_{parcel on sfc} = \eta$ 

 $\therefore \quad w_{\text{parcel on sfc}} = \frac{D}{Dt} z_{\text{parcel on sfc}} = \frac{D\eta}{Dt} = \frac{\partial\eta}{\partial t} + u \frac{\partial\eta}{\partial x}$  $\therefore \quad \frac{\partial\phi}{\partial z} = \frac{\partial\eta}{\partial t} + u \frac{\partial\eta}{\partial x} \quad \text{on } z = \eta(x,t)$ 



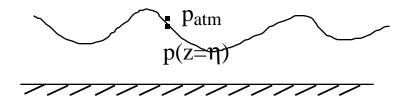
For infinitesimal waves, neglect products of small quantities (nonlinear terms). Neglect  $u \frac{\partial \eta}{\partial x}$ .

 $\therefore \frac{\partial \phi}{\partial z} = \frac{\partial \eta}{\partial t} \text{ on } z = \eta(x,t) \text{ [still nonlinear: } \eta \text{ is affected by flow]}$ Exapnd l.h.s. in a Taylor series about z = 0:

$$\frac{\partial \phi}{\partial z}\Big|_{z=\eta} = \frac{\partial \phi}{\partial z}\Big|_{z=0} + \frac{\partial}{\partial z}\left(\frac{\partial \phi}{\partial z}\right)\Big|_{z=0}\eta + \dots$$
[\_\_\_\_\_\_\_\_]
neglect these terms since  $\eta$  is small
$$\frac{z=0}{z=\eta}$$

 $\therefore \quad \frac{\partial \phi}{\partial z} = \frac{\partial \eta}{\partial t} \text{ on } z = 0 \quad \underline{\text{Linearized kinematic b.c. on free sfc.}}$ [Neglected a nonlinear term <u>and put b.c. at z=0 instead of z=\eta</u>]

<u>Dynamic b.c.</u> on free surface: pressure is <u>continuous</u> across air/sea interface. So pressure on liquid side of interface = atm pressure.



 $p(z = \eta(x,t)) = p_{atm}$ 

Translate this b.c. into a b.c. on  $\phi$ , using Bernoulli's eq<sup>n</sup> for unsteady, irrot flow:

 $\frac{\partial \phi}{\partial t} + \frac{q^2}{2} + \frac{p}{\rho} + gz = C \quad (\text{same const everywhere})$ Apply it on free sfc z= $\eta$  (liquid side), w/ p=p<sub>atm</sub>, and neglect q<sup>2</sup>:

$$\therefore \frac{\partial \phi}{\partial t} + \frac{p_{atm}}{\rho} + g\eta = C$$
  
$$\therefore \frac{\partial \phi}{\partial t} + g\eta = (C - \frac{p_{atm}}{\rho}) = \text{const}$$

Set const = 0. [Flow doesn't care about it. Equivalently, define  $\phi_{new} = \phi_{old} + \text{const t}$ , and plug into above eqn. The const cancels out, but since  $\vec{u} = \nabla \phi$  and  $\nabla \phi_{new} = \nabla \phi_{old}$ , the flow is unchanged.]

$$\therefore \quad \underbrace{\frac{\partial \phi}{\partial t}}_{\downarrow} + g\eta = 0 \quad \text{at } z = \eta$$

Expand it in a Taylor series in z about z = 0:

[neglect since  $\eta$  is small]

$$\therefore \quad \frac{\partial \phi}{\partial t}\Big|_{z=\eta} = \left.\frac{\partial \phi}{\partial t}\right|_{z=0} + \left.\frac{\partial}{\partial z}\left(\frac{\partial \phi}{\partial t}\right)\right|_{z=0} \eta + \text{h.o.t.}$$

: <u>linearized dynamic</u> free sfc b.c. is:

$$\frac{\partial \phi}{\partial t} + g\eta = 0$$
 at  $z = 0$ 

Want to solve  $\frac{\partial^2 \phi}{\partial x^2} + \frac{\partial^2 \phi}{\partial z^2} = 0$  subject to the above b.c.

Consider a "wavy" pattern for  $\eta$ :  $\eta = a \cos(kx - \omega t)$ 

From either top b.c., suspect  $\phi \propto \sin(kx - \omega t) w/a z$  dependence.

Trial solution:  $\phi = f(z) \sin(kx - \omega t)$  [will it work?] Plug it into Laplace's eq<sup>n</sup>, get:

$$-k^{2} f \sin(kx - \omega t) + \frac{d^{2} f}{dz^{2}} \sin(kx - \omega t) = 0$$

 $\therefore \frac{d^2f}{dz^2} - k^2f = 0 \quad a 2^{nd} \text{ order linear const coeff ode}$ 

Trial soln for f:  $f = e^{mz}$ . Plug into ode, get:

$$m^2 - k^2 = 0$$
,  $\therefore m = k \text{ or } - k$ 

So general sol<sup>n</sup> for f is:  $f = A e^{kz} + B e^{-kz}$ and so  $\phi$  becomes:

(\*) 
$$\phi = (A e^{kz} + B e^{-kz}) \sin(kx - \omega t)$$

Now use b.c. to pin down A, B.

Lower kinematic b.c. is

$$\frac{\partial \phi}{\partial z} = 0$$
 at  $z = -H$ .

Apply it in (\*), get

$$\left(A k e^{kz} - k B e^{-kz}\right) \bigg|_{z=-H} \sin(kx - \omega t) = 0$$
  

$$\therefore A k e^{-kH} - k B e^{kH} = 0$$
  

$$\therefore A = B e^{2kH}$$

Top linearized kinematic b.c. is

$$\frac{\partial \phi}{\partial z} = \frac{\partial \eta}{\partial t}$$
 on  $z = 0$ .

Apply it in (\*), get

$$\therefore \left( Ake^{kz} - kBe^{-kz} \right) \bigg|_{z=0} \sin(kx - \omega t) = -a(-\omega) \sin(kx - \omega t)$$

$$\therefore k(A - B) = a\omega$$

2 linear algebraic eqns for 2 unknowns, A, B. Solution is:

$$A = \frac{a\omega}{k(e^{2kH} - 1)} e^{2kH}, \quad B = \frac{a\omega}{k(e^{2kH} - 1)}$$
  
$$\therefore \quad \phi = \frac{a\omega}{k(e^{2kH} - 1)} (e^{2kH} e^{kz} + e^{-kz}) \sin(kx - \omega t)$$

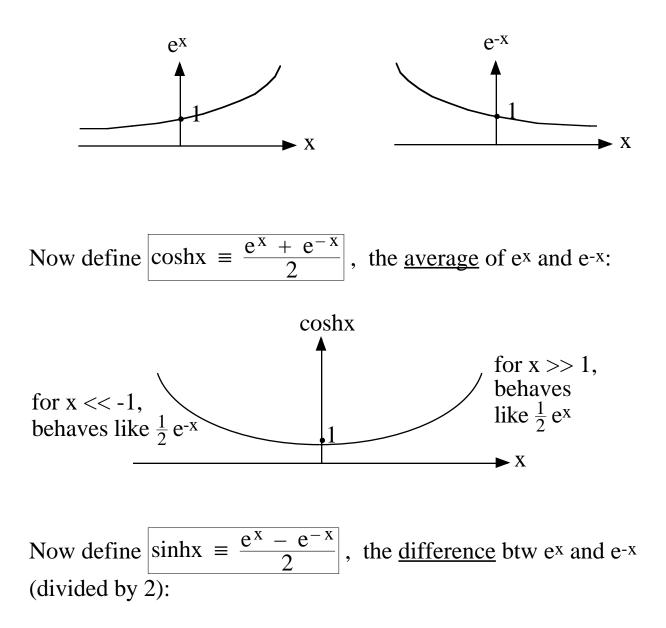
factor out  $e^{kH}$  from top and bottom, and mult by  $\frac{2}{2}$ 

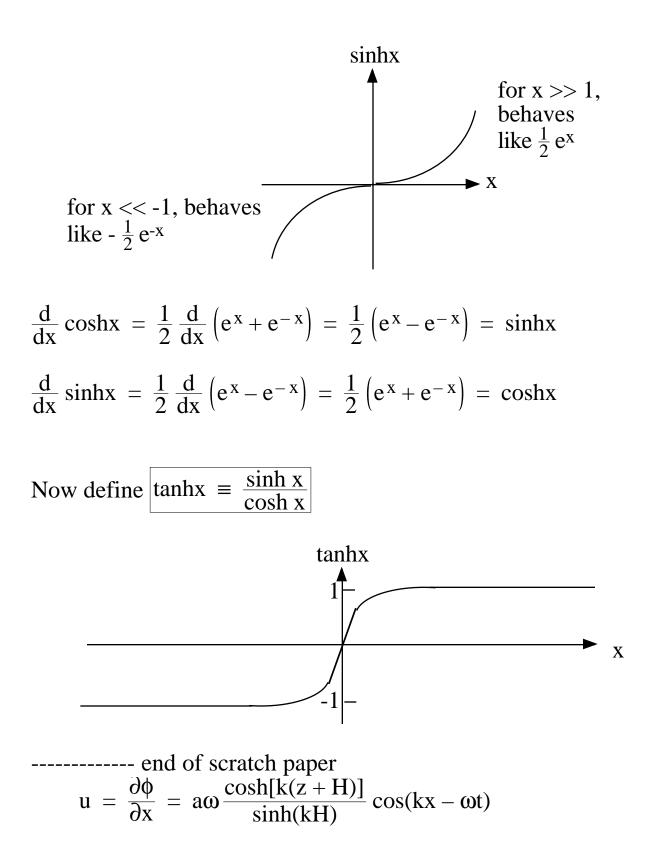
$$\therefore \quad \phi = \frac{a\omega}{k} \frac{2}{e^{kH} - e^{-kH}} \left( \frac{e^{k(z+H)} + e^{-k(z+H)}}{2} \right) \sin(kx - \omega t)$$

$$\therefore \quad \phi = \frac{a\omega}{k} \frac{\cosh\left[k\left(z+H\right)\right]}{\sinh\left(kH\right)} \sin(kx - \omega t)$$

----- scratch paper, review of hyperbolic functions

here's how e<sup>x</sup> and e<sup>-x</sup> behave:





w = 
$$\frac{\partial \phi}{\partial z}$$
 =  $a\omega \frac{\sinh[k(z+H)]}{\sinh(kH)} \sin(kx - \omega t)$ 

Get p from linearized Bernoulli eqn:

$$\frac{p}{\rho} = \text{const} - gz - \frac{\partial \phi}{\partial t}$$
$$= \text{const} - gz + \frac{a\omega^2}{k} \frac{\cosh[k(z+H)]}{\sinh(kH)} \cos(kx - \omega t)$$

Now apply linearized dynamic b.c.:  $\frac{\partial \phi}{\partial t} + g\eta = 0$  on z = 0

$$\therefore -\frac{a\omega^2}{k}\frac{\cosh(kH)}{\sinh(kH)}\cos(kx-\omega t) + g a \cos(kx-\omega t) = 0$$

$$\omega^2 = gk \tanh(kH)$$

$$\therefore \quad \overline{\omega = \sqrt{gk \tanh(kH)}} \quad \underline{dispersion \ relation} \ [\omega = \omega(k)]$$
$$c = \frac{\omega}{k}, \qquad \therefore \ c = \sqrt{\frac{g}{k} \tanh(kH)}$$

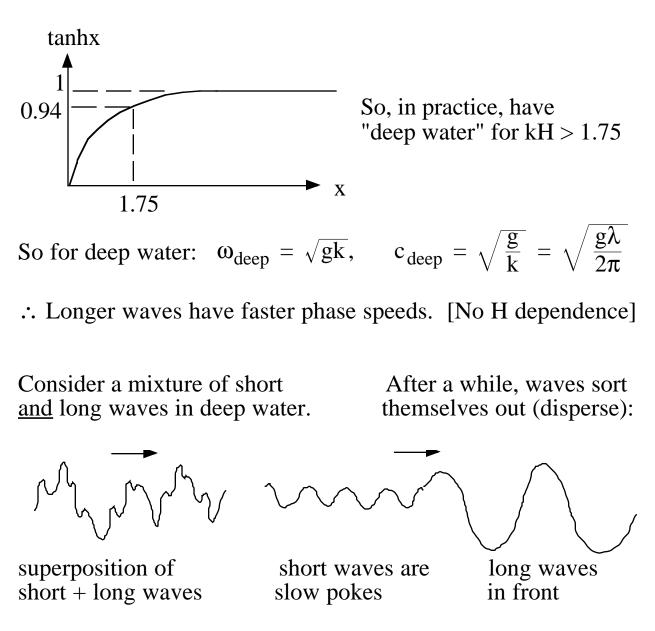
Note: c,  $\omega$  indep of amplitude a (a feature of linear waves).

Now consider the limiting cases of "deep-water" and "shallowwater" surface gravity waves.

"deep-water" condition:  $\lambda \ll H$  or  $\frac{H}{\lambda} \gg 1$  or  $kH \gg 1$ 

 $\therefore \tanh(kH) \approx 1$ 

Actually,  $tanh(kH) \approx 1$  even for  $H/\lambda$  not too big:



When there's a storm way out at sea, the first waves to reach the shore are the long waves (low frequency waves). "shallow-water" condition:  $\lambda \gg H$  or  $\frac{H}{\lambda} \ll 1$  or kH  $\ll 1$ 

$$\therefore \tanh(kH) = \frac{e^{kH} - e^{-kH}}{e^{kH} + e^{-kH}} = \frac{(1 + kH + ...) - (1 - kH + ...)}{(1 + kH + ...) + (1 - kH + ...)}$$

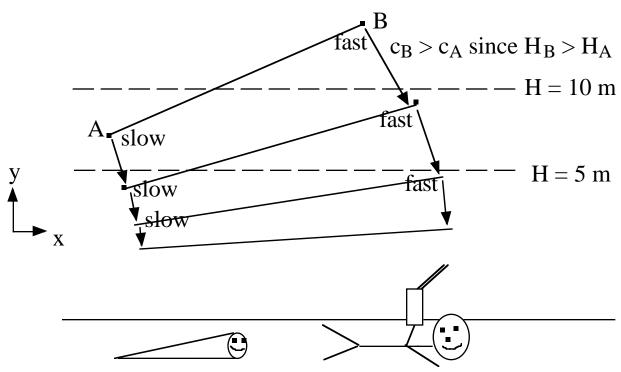
$$= \frac{2kH + ...}{2 + ...} \approx kH$$
  

$$\therefore \omega_{\text{shallow}} = \sqrt{gk \ kH} = k \ \sqrt{gH}$$
  

$$\therefore c_{\text{shallow}} = \frac{\omega_{\text{shallow}}}{k} = \sqrt{gH} \quad \text{indep of } k!$$

So shallow-water waves are <u>non-dispersive</u>. For shallow water waves,  $c \downarrow as H \downarrow$ 

Consider wave crests approaching a beach obliquely:



Wave crests turn as they approach beach, end up || to beach. A case of <u>wave refraction</u> (bending of wave fronts in inhomogeneous media -- in this case variable H(x)).

recall that for general surface gravity waves:

$$\omega = \sqrt{gk \tanh(kH)}$$
,  $c = \sqrt{\frac{g}{k} \tanh(kH)}$ 

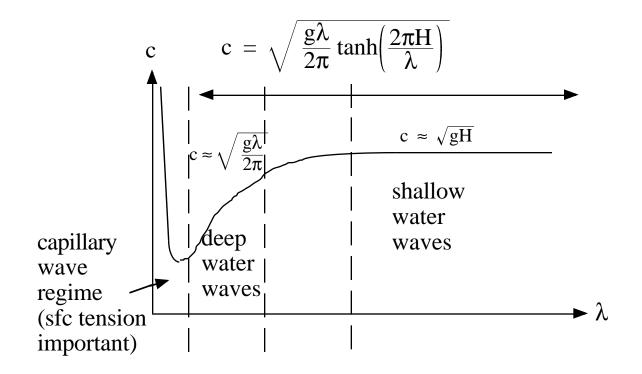
and that for "shallow water":  $kH \ll 1$  so we're led to

$$\omega_{\text{shallow}} = k \sqrt{gH}$$
,  $c_{\text{shallow}} = \sqrt{gH}$ 

Now consider pressure in shallow water conditions:

$$\frac{p_{\text{shallow}}}{\rho} = \text{const} - gz + \frac{a\omega^2}{k} \frac{\cos[k(z+H)]}{\sinh(kH)} \cos(kx - \omega t)$$
$$= \text{const} - gz + ag\cos(kx - \omega t)$$
$$= \text{const} - g(z - \eta) \quad \text{Hydrostatic pressure distribution.}$$

Now look at phase speed c for the general surface wave case (deep/shallow/whatever):



"deep" or "shallow" depends on  $\lambda$  relative to H. Water that's 100 m deep is "deep" for  $\lambda = 10m$  but "shallow" for  $\lambda = 1000$  m.

Derive <u>streamfunction</u> for surface gravity wave [recall this is a 2D incomp flow]

$$\frac{\partial \Psi}{\partial z} = u = \frac{\partial \Phi}{\partial x} = a\omega \frac{\cosh[k(z+H)]}{\sinh(kH)}\cos(kx-\omega t)$$

integrate w.r.t. z:

(1) 
$$\psi = \frac{a\omega}{k} \frac{\sinh[k(z+H)]}{\sinh(kH)} \cos(kx - \omega t) + F(x,t)$$

Similarly,  $\frac{\partial \Psi}{\partial x} = -w = ...$  Integrate w.r.t. x, to get:

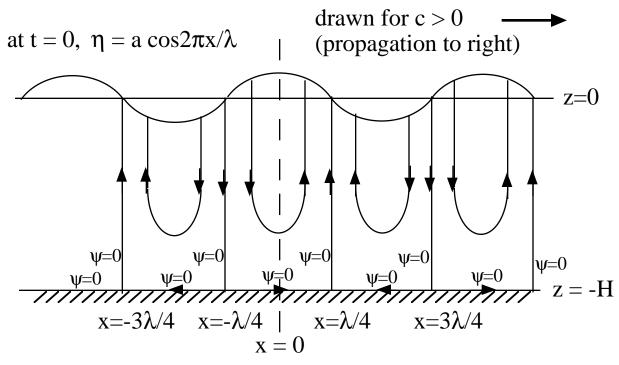
(2) 
$$\Psi = \frac{a\omega}{k} \frac{\sinh[k(z+H)]}{\sinh(kH)} \cos(kx - \omega t) + G(z,t)$$

F(x,t) = G(z,t) but f<sup>n</sup> of x can't be a f<sup>n</sup> of z -- so no x or z dependence. So F(x,t) = G(z,t) = E(t). But E(t) is <u>irrelevant</u> since u, w only care about spatial derives of  $\psi$ . So take E(t) = 0.

$$\therefore \quad \Psi = \frac{a\omega}{k} \frac{\sinh[k(z+H)]}{\sinh(kH)} \cos(kx - \omega t)$$

Graph streamlines ( $\psi = \text{const}$ ) at t = 0.

$$\psi = 0$$
 for:  $z = -H$  and for:  $k = \pm \frac{\pi}{2}, \pm \frac{3\pi}{2}, \pm \frac{5\pi}{2}, \dots$   
 $2\pi/\lambda$   
 $x = \pm \frac{\lambda}{4}, \pm \frac{3\lambda}{4}, \pm \frac{5\lambda}{4}, \dots$ 



[Get dirn of flow (arrows) from soln for u or w, or consider: for pattern moving toward right,  $\eta$  is rising to right of crest (so w>0 there) and  $\eta$  is falling to left of crest (so w<0 there). Arrows on

bottom give sense of horiz conv/div needed to support this w field. Clearly parcel velocity differs from phase speed.].

## Group Velocity

Concept of group velocity is appropriate for many different types of waves (not just surface gravity waves).

Consider <u>2 waves</u> of <u>equal amplitude</u> and <u>slightly different</u> <u>frequency and wavelength</u> moving in <u>same</u> direction:

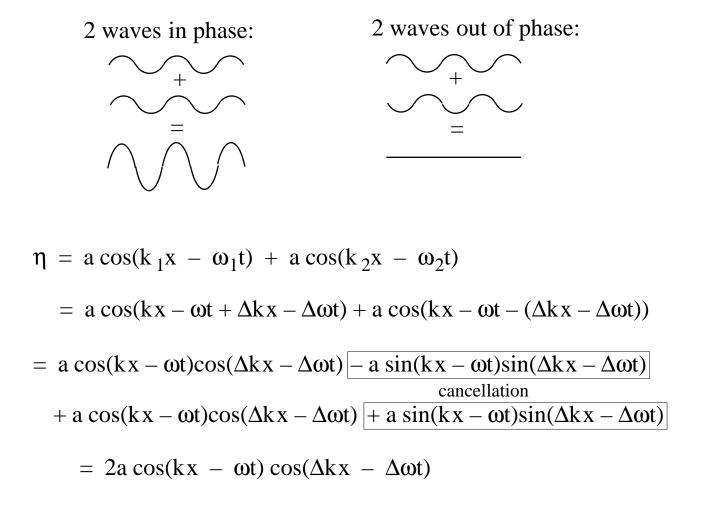
- $\omega_1 = \omega + \Delta \omega, \qquad k_1 = k + \Delta k$
- $\omega_2 = \omega \Delta \omega, \qquad k_2 = k \Delta k$
- assume  $\frac{\Delta\omega}{\omega} \ll 1$ ,  $\frac{\Delta k}{k} \ll 1$

Because of dispersion relation,  $\Delta \omega$  is related to  $\Delta k$ .

Mean frequency is: 
$$\frac{\omega_1 + \omega_2}{2} = \frac{\omega + \Delta \omega + \omega - \Delta \omega}{2} = \omega$$

Mean wavenumber is:  $\frac{k_1 + k_2}{2} = \dots = k$ 

Where the waves are in phase (or nearly so) they combine to form a wave of twice amplitude. Where they're out of phase, they kill each other off.



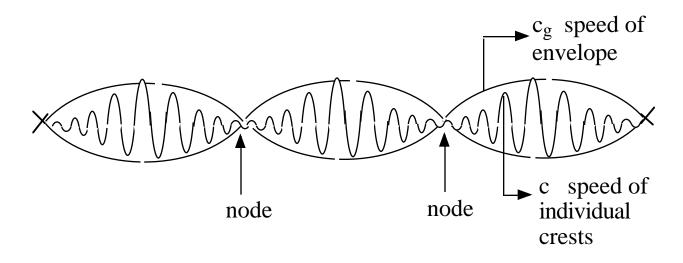
$$\therefore \eta = A \underbrace{\cos(kx - \omega t)}_{\downarrow} \quad \text{where } A \equiv 2a \cos(\Delta kx - \Delta \omega t)$$
  
effective carrier wave (mean wave)  
amplitude

Effective amplitude A is itself a wave with wavelength  $\lambda_{\text{amplitude}} = \frac{2\pi}{\Delta k} \implies \frac{2\pi}{k} = \lambda_{\text{carrier wave}}$ A propagates at speed  $\frac{\Delta \omega}{\Delta k}$  where  $\Delta \omega$  is related to  $\Delta k$  by dispersion rel<sup>n</sup>. For small  $\Delta k$ ,  $\frac{\Delta \omega}{\Delta k} \rightarrow \frac{d\omega}{dk}$ . Define  $c_g \equiv \frac{d\omega}{dk}$ 

or 
$$\vec{c}_g \equiv \frac{d\omega}{dk} \hat{i}$$
 Group velocity. A vector.

c is phase speed of crests [not a vector, see fig. 7.3 Kundu)

 $c_g$  is speed of <u>envelope</u> of crests.



Energy is trapped between nodes  $\therefore$  energy propagates at speed of nodes (speed of envelope), i.e. speed  $c_g$ , not phase speed c.

For <u>deep-water sfc waves</u>:

$$\omega = \sqrt{gk}$$
$$c = \sqrt{\frac{g}{k}}$$
$$c_g = \frac{d\omega}{dk} = \frac{1}{2\sqrt{k}}\sqrt{g}$$

$$\therefore \quad \boxed{c_g = \frac{1}{2}c} \text{ [since } c > c_g, \text{ individual crests move through} \\ \text{envelope, die at nodes]}$$

For shallow water sfc waves:

$$\omega = k \sqrt{gH}$$

$$c = \sqrt{gH}$$

$$c_g = \frac{d\omega}{dk} = \sqrt{gH}$$

$$\therefore \quad c_g = c$$